

SULFUR SPECIATION IN THE MARTIAN REGOLITH COMPONENT IN SHERGOTTITE GLASSES. M. N. Rao¹,

L. E. Nyquist², S. Sutton³, and J. Huth⁴. ¹Jacobs-ESCG, Johnson Space Center, Houston. TX. E-mail: nageswara.rao@nasa.gov.

²ARES,NASA,Johnson Space Center, Houston. TX. ³Department of Geological Sciences, University of Chicago, Chicago.IL.

⁴Max-Planck-Institute fuer Chemie, Saarstrasse 23, Mainz. Germany.

Introduction: We have shown that Gas-Rich Impact-Melt (GRIM) glasses in Shergotty, Zagami, and EET79001 (Lith A and Lith B) contain Martian regolith components that were molten during impact and quenched into glasses in voids of host rock materials based on neutron-capture isotopes, i.e., ¹⁵⁰Sm excesses and ¹⁴⁹Sm deficits in Sm, and ⁸⁰Kr excesses produced from Br [1, 2]. These GRIM glasses are rich in S-bearing secondary minerals [3,4]. Evidence for the occurrence of CaSO₄ and S-rich aluminosilicates in these glasses is provided by CaO-SO₃ and Al₂O₃ – SO₃ correlations, which are consistent with the finding of gypsum laths protruding from the molten glass in EET79001 (Lith A) [5]. However, in the case of GRIM glasses from EET79001 (Lith B), Shergotty and Zagami, we find a different set of secondary minerals that show a FeO-SO₃ correlation (but no MgO-SO₃ correlation), instead of CaO-SO₃ and Al₂O₃-SO₃ correlations observed in Lith A. These results might indicate different fluid-rock interactions near the shergottite source region on Mars. The speciation of sulfur in these salt assemblages was earlier studied by us using XANES techniques [6], where we found that Lith B predominantly contains Fe-sulfide globules (with some sulfate). On the other hand, Lith A showed predominantly Ca-sulfite/sulfate with some FeS. Furthermore, we found Fe to be present as Fe²⁺ indicating little oxidation, if any, in these glasses.

To examine the sulfide-sulfate association in these glasses, we studied their Fe/Ni ratios with a view to find diagnostic clues for the source fluid. The Fe-sulfide mineral (Fe_{0.93}Ni_{0.3}S) in EET79001, Lith A is pyrrhotite [7, 8]. It yields an Fe/Ni ratio of 31. In Shergotty, pyrrhotite occurs with a molar ratio of Fe:S of 0.94 and a Ni abundance of 0.12% yielding a Fe/Ni ratio of ~500 [8]. In this study, we determined a NiO content of ~0.1% and FeO/NiO ratio of ~420 in S-rich globules in #507 (EET79001, Lith B) sample using FE-SEM. In the same sample (bulk), using EMPA, we determined a FeO/NiO ratio of ~700 (raster mode). Using similar techniques, we determined a NiO content of ~0.015% and a FeO/NiO ratio of ~800 in #506 (EET79001, Lith A). Moreover, a NiO content of ~150 ppm and 6.1% FeO were found in Lith A GRIM glasses using neutron activation analysis [9] yielding a FeO/NiO ratio of ~420. The FeO/NiO ratios in secondary mineral phases in S-rich pockets of EET79001 (Lith A/B) and Shergotty are high (~400) compared to the FeO/NiO ratio of 31 in Lith A pyrrhotite. These results suggest similar kind of fluids interacted with different rock materials to yield the observed variations in GRIM glasses in EET79001 Lith A and B.

References: [1] Rao M. N. et al. 2008. *LPSC 40th*, #1361. [2] Rao M. N. et al. 2002. *Icarus*, 156:352-372. [3] Rao M. N. et al. 1999. *GRL*, 26:3265-3268. [4] Gooding J. L. et al. 1988. *GCA*, 52:909-915. [5] Rao M. N. and McKay D. S. 2003. 6th, *Inter. Mars Conf. J.P.L.* Pasadena, CA. #3130. [6] Sutton S.R. et al. 2008, *XXXIX LPSC*, #1961. [7] McSween H. Y. and Jarosewich E. 1983. *GCA*, 47: 1501-1513. [8] Smith J.V. and Hervig R. L. 1979. *Meteoritics*, 14: 121-142. [9] Smith M. R. et al. 1984. *LPSC 14 (JGR)*, 89: B612-630.